

Water Waves

no viscosity.

Euler eqs : water inviscid, incompressible fluid.

use my notes on WWW instead of going through details on the board

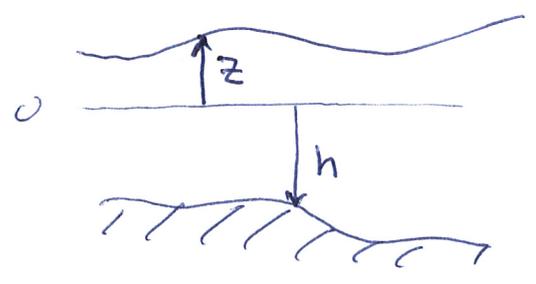
Euler eqs.

$$\begin{cases} u_x + v_y + w_z = 0 \\ u_t + u u_x + v u_y + w u_z = -\frac{1}{\rho} p_x \\ v_t + u v_x + v v_y + w v_z = -\frac{1}{\rho} p_y \\ w_t + u w_x + v w_y + w w_z = -\frac{1}{\rho} p_z - g \end{cases}$$

g : gravity, ρ : density, $\vec{P} = (P_x, P_y, P_z)$: pressure

$\vec{U} = (u, v, w)$ fluid velocity.

let $z = \eta(x, y, t)$: free surface (top)
 $z = -h(x, y)$: bottom topography:



[S, p 53-69]
 → photocopy to students.

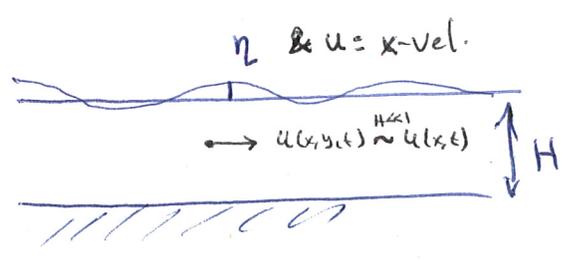
Boussinesq eq + KdV eq

- consider irrotational flow
- 2D
- shallow, flat bottom
- low order nonlinearities:
- dimensionalization

$$\Rightarrow \begin{cases} \eta_t + [(1 + \delta \eta) u]_x - \frac{\delta}{6} u_{xxx} = 0 \\ u_t + \delta u u_x + \eta_x - \frac{\delta}{2} u_{xxt} = 0 \end{cases}$$

Boussinesq eqns. (1)

$\delta = (\frac{H}{L})^2 \ll 1$, L : typical wave length.



Linearize eqs:

$$u = u^e + \epsilon u^p$$

$$\eta = \eta^e + \epsilon \eta^p$$

$$\begin{matrix} u^e = 0 \\ \eta^e = 0 \end{matrix}$$

steady state solutions. $|\epsilon| \ll 1$ perturbation param.

Euler Eq. \Rightarrow

$$\begin{cases} \epsilon \eta_{tt}^p + [(1 + \delta \epsilon \eta^p) \epsilon u_x^p]_x - \frac{\delta}{6} \epsilon u_{xxx}^p = 0 \\ \epsilon u_{tt}^p + \epsilon^2 u^e u_x^p + \epsilon \eta_x^p - \frac{\delta}{2} \epsilon u_{xxt}^p = 0 \end{cases}$$

Suppose u & η variations very soft so that we neglect higher order derivatives: Keeping $O(\epsilon)$ terms:

$$\begin{cases} \eta_{tt}^p + u_x^p = 0 \\ u_{tt}^p + \eta_x^p = 0 \end{cases} \quad \left\| \begin{array}{l} \text{equivalent to } \delta = 0 \\ \text{or } u^p = \eta^p \end{array} \right.$$

$$\Rightarrow \begin{cases} \eta_{tt}^p + u_x^p = 0 \\ u_{tt}^p + \eta_x^p = 0 \end{cases} \quad \ominus \quad \boxed{\eta_{tt}^p - \eta_{xx}^p = 0} \quad \text{②}$$

linear wave eq.

[and $u^p = \eta^p$ is a solution in the linear case.]

Similar structure for u^p . Since wave in η and u should be related ~~but not in the linear case~~ $\Rightarrow u^p = \eta^p$

dropped u^p from wave eq.

let us take $u = \eta(x,t) + \delta A(x,t) + O(\delta^2)$ ③ and determine $A(x,t)$ such that ③ is sol. to ①

$$\text{③ in ①} : \begin{cases} \eta_t + [(1 + \delta \eta)(\eta + \delta A)]_x - \frac{\delta}{6} [\eta + \delta A]_{xxx} = 0 \\ (\eta + \delta A)_{tt} + \delta (\eta + \delta A)(\eta + \delta A)_x + \eta_x + \frac{\delta}{2} [\eta + \delta A]_{xxt} = 0 \end{cases}$$

now keep only terms $< O(\delta^2)$

$$\Rightarrow \begin{cases} \eta_t + [\eta + \delta \eta^2 + \delta A]_x - \frac{\delta}{6} \eta_{xxx} = 0 \\ \eta_{tt} + \delta A_{tt} + \delta \eta \eta_x + \eta_x + \frac{\delta}{2} \eta_{xxt} = 0 \end{cases}$$

$$\Rightarrow \begin{cases} \eta_t + \eta_x + 2\delta \eta \eta_x + \delta A_x - \frac{\delta}{6} \eta_{xxx} = 0 \\ \eta_{tt} + \delta A_{tt} + \delta \eta \eta_x + \eta_x - \frac{\delta}{2} \eta_{xxt} = 0 \end{cases}$$

Now consider only wave traveling to the right i.e. $\eta_t + \eta_x = 0$

$$\Rightarrow \eta_t = -\eta_x \Rightarrow \cancel{\eta_t} = -\cancel{\eta_x} \quad \delta \eta_{xxt} = -\delta \eta_{xxx}$$

$$\begin{cases} \eta_t + \eta_x + 2\delta \eta \eta_x + \delta A_x - \frac{\delta}{6} \eta_{xxx} = 0 & (4) \\ 2\eta_t + \eta_x + \delta(\eta \eta_x) + \delta A_t + \frac{\delta}{2} \eta_{xxx} = 0 \end{cases}$$

If the 2 eqns are to be consistent (ie solve for A) ← solvability condition.

$$(4a) - (4b) \Rightarrow \delta \eta \eta_x + \delta(A_x - A_t) - \frac{2\delta}{3} \eta_{xxx} = 0$$

$$A_x - A_t = -\eta \eta_x + \frac{\delta^2}{3} \eta_{xxx}$$

If $A = -\frac{\eta^2}{4} + \frac{\eta_{xx}}{3}$ (5)

$$A_x - A_t = -\frac{2\eta \eta_x}{4} + \frac{\eta_{xxx}}{3} + \frac{2\eta \eta_t}{4} - \frac{\eta_{xxt}}{3}$$

$$\begin{aligned} \eta_t = -\eta_x \\ &= -\frac{1}{2} \eta \eta_x + \frac{\eta_{xxx}}{3} + \frac{1}{2} \eta \eta_x + \frac{\delta}{3} \eta_{xxx} \end{aligned}$$

Since $A_x = -A_t \Rightarrow \frac{d}{dx} \left(-\frac{\eta^2}{4} + \frac{\eta_{xx}}{3} \right) = -\frac{d}{dt} \left(-\frac{\eta^2}{4} + \frac{\eta_{xx}}{3} \right)$
 $\therefore A = \frac{1}{2} \left(-\frac{1}{2} \eta^2 + \frac{2}{3} \eta_{xx} \right)$ should work.

$$= -\eta \eta_x + \frac{2}{3} \eta_{xxx} \checkmark$$

\therefore use (4a) or (4b) with (5):

$$(5) \text{ in } (4a) \quad \eta_t + \eta_x + 2\delta \eta \eta_x + \delta \left(-\frac{2\eta \eta_x}{4} + \frac{\eta_{xxx}}{3} \right) - \frac{\delta}{6} \eta_{xxx} = 0$$

$$\Rightarrow \eta_t + \eta_x + (2 - \frac{1}{2}) \delta \eta \eta_x + \delta \left(\frac{1}{3} - \frac{1}{6} \right) \eta_{xxx} = 0$$

$$\Rightarrow \boxed{\eta_t + \eta_x + \frac{3}{2} \delta \eta \eta_x + \frac{\delta}{6} \eta_{xxx} = 0} \quad (6)$$

Now transform $t \rightarrow \underbrace{t}_{\text{rescale time}}, x \rightarrow \underbrace{x - (1-\delta)t}_{\text{co-moving reference frame @ velocity } (1-\delta)}$

HW

$$\Rightarrow \boxed{\eta_t + \eta_x + \frac{3}{2} \eta \eta_x + \frac{1}{6} \eta_{xxx} = 0} \quad (7) \quad \text{Korteweg-de Vries Eq. (KdV)}$$



↑ model for shallow water wave equations in one direction, no dissipation and lowest order dispersion (η_{xxx})

Elementary solutions of KdV

$$U_t + C(u)U_x = 0 \quad \leftarrow \text{vel.}$$

18
(12)

Under, yet another transformation, KdV (7) can be rewritten as HK

$$U_t + 6UU_x + U_{xxx} = 0 \quad (8) \text{ KdV [Drazin]}$$

→ no parameters: transf. eliminated all params.
 → Drazin has \ominus but this is just $u \rightarrow -u$

Scale invariance:

transf: $X = k^a x \Rightarrow \partial_x = k^a \partial_x$
 $T = k^b t \Rightarrow \partial_T = k^b \partial_t$
 $U = k^c u$

$$\text{KdV} \Rightarrow k^{b-c} U_T - 6 k^{a-2c} U U_x + k^{+3a-c} U_{xxx} = 0$$

$$\begin{aligned} b-c = a-2c = 3a-c &\Rightarrow a=1, b=3, c=-2 \quad \checkmark \\ 3+2 = 1+4 = 3+2 & \end{aligned}$$

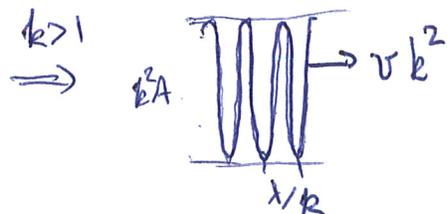
\therefore if $X = kx, T = k^3 t, U = u/k^2 \Rightarrow \text{KdV} \Leftrightarrow U_T + 6UU_x + U_{xxx} = 0$
 i.e. the solution is the same!

\therefore if $u = A f(x-ut)$ is sol to KdV
 then $U = A f(kx-ut)$ is also a sol. to KdV

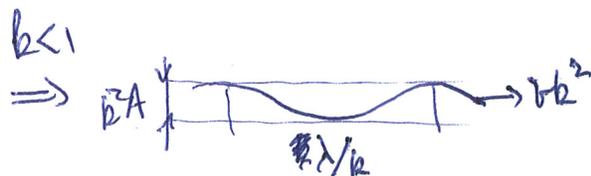
$$\Rightarrow \frac{u}{k^2} = U = A f(kx - ut)$$

$$\begin{aligned} \Rightarrow u(x,t) &= k^2 A f(k(x - ut)) \\ &= k^2 A f(k(x - ut)) \end{aligned}$$

$$v' = k^2 v$$



∇ This is true $\forall f(x)$!



Galilean invariance

If $f(x-vt)$ is a sol $\Rightarrow \lambda + f(x-\bar{v}t)$ is also a solution $[v = \bar{v} - 6\lambda]$ HW

\therefore one can add/subtract water to canal and the slope remains the same and only speed changes. do NOT give. Just say: $v = \bar{v} - f(x)$

Traveling solutions: let us look for traveling solutions:

$$u(x,t) = f(x-ct) = f(\zeta), \quad \zeta = x-ct$$

$$\text{KdV} \Rightarrow \boxed{-cf' + 6ff' + f''' = 0} \quad (9) \quad (\cdot)' \equiv \frac{\partial}{\partial \zeta}(\cdot)$$

\leftarrow 1 parameter family of ODEs

\therefore any solution to ODE (9) is a sol. to KdV.

Solitary wave: integrate (9) using $f(\pm\infty) = 0 = f' = f''$

$$\int dz \Rightarrow -c f + 3f^2 + f'' = A \leftarrow \text{constant.}$$

but since we require $f(\pm\infty) = 0 \Rightarrow A = 0$

$$\Rightarrow -c f + 3f^2 + f'' = 0$$

$$\times f' \Rightarrow -c f f' + 3 f^2 f' + f'' f' = 0$$

$$\int dz \Rightarrow -\frac{c}{2} f^2 + f^3 + \frac{1}{2} (f')^2 = B$$

Again BCs $\Rightarrow B = 0 \Rightarrow (f')^2 = c f^2 - 2 f^3 = f(c - 2f)$

$$\Rightarrow f' = \pm f \sqrt{c-2f}$$

$$\Rightarrow \frac{df}{\sqrt{c-2f}} = \pm f \sqrt{c-2f} \Rightarrow \int \frac{df}{\sqrt{c-2f}} = \pm \int dz$$

$$\Rightarrow \pm \frac{2}{\sqrt{c}} \tanh^{-1} \left[\frac{\sqrt{c-2f}}{\sqrt{c}} \right] = \zeta - x_0$$

$$\Rightarrow \pm \sqrt{c-2f} = \sqrt{c} \tanh \frac{\sqrt{c}}{2} (\zeta - x_0)$$

$$\Rightarrow 2f = -c \tanh^2(\dots) + c = c \operatorname{sech}^2(\dots)$$

$$\Rightarrow f(x-ct) = \frac{c}{2} \operatorname{sech}^2 \left[\frac{\sqrt{c}}{2} (\zeta - x_0) \right]$$

$$\text{or } u(x,t) = \frac{c}{2} \operatorname{sech}^2 \left[\frac{\sqrt{c}}{2} (x-ct-x_0) \right]$$

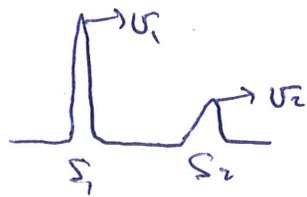
better use KdV.m

Mathlab PDE integrator

solitary solution to KdV. scale invariance

Superposition?

→ No, eq. is nonlinear but if we start 2 sech^2 solitons very far away it is almost an exact sol:



→ Q: what happens as $t \uparrow$?

- superposition?
- phase shift.

KdV.m
2 sol. IC.

→ Name SOLITON

⚠ KdV is the simplest eq. that incorporates nonlinearity & dispersion: i.e. it is a GENERIC eq. that will often show up in nonlinear propagation models.

- Internal gravity waves, stratified fluids
- waves in rotating atmosphere
- ion-acoustic waves in plasmas
- pressure waves in liquid-gas bubble mixtures
- nonlinear (anharmonic) lattices
- thermal packets in nonlinear crystals
- ...

General waves of permanent form in KdV

KdV: $u_t + 6uu_x + u_{xxx} = 0$

traveling wave: $u(x,t) = f(\xi)$ $\xi = x - ct$

⇒ $-cf' + 6ff' + f''' = 0$

$\int d\xi = 1$ $-cf + 3f^2 + f'' = -A/2$

⇒ $f'' = cf - 3f^2 + A/2$

⇒ $2f'' = 2cf - 6f^2 + A$

(← now we don't require $u(\pm\infty) = u_x(\pm\infty) = 0$)
A to be determined by BCs.

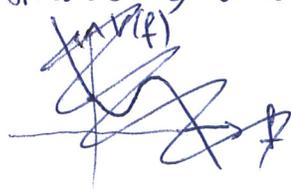
Remember Newton's law: $m\ddot{y} = F(y) = -\frac{dV}{dy}$ ← potential.

Here if potential $V(f) = -cf^2 + 2f^3 + Af + B$ ($m=2$)

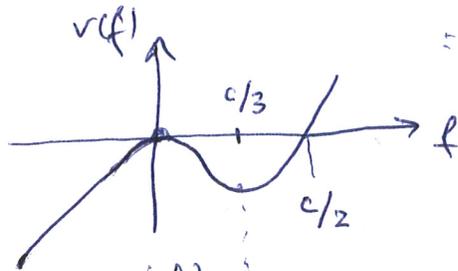
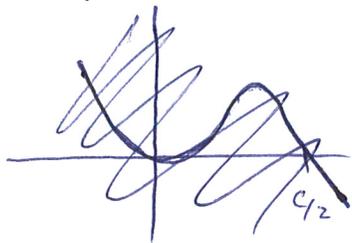
⇒ $m f'' = -\frac{dV}{df}$

∴ any possible orbit of $m\ddot{y} = -\frac{dV}{dy}$ under influence of V is a valid traveling wave of KdV!

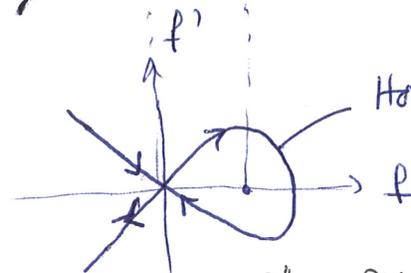
Plot potential:



For our previous sech soliton: $A=B=0 \Rightarrow v(f) = 2f^3 - cf^2 = f^2(2f-c)$ 21
(2)

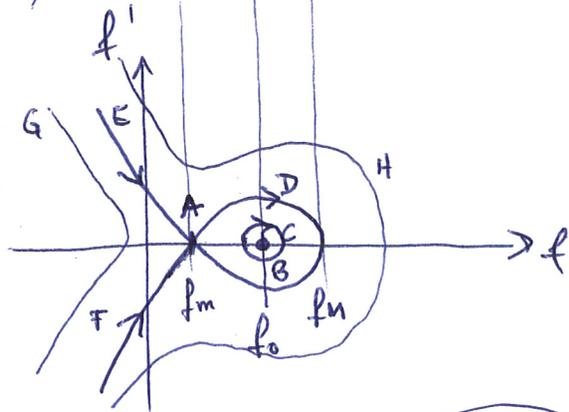
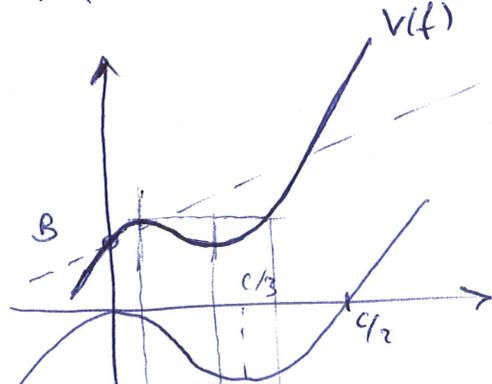


Phase plane for 2nd-order ODE

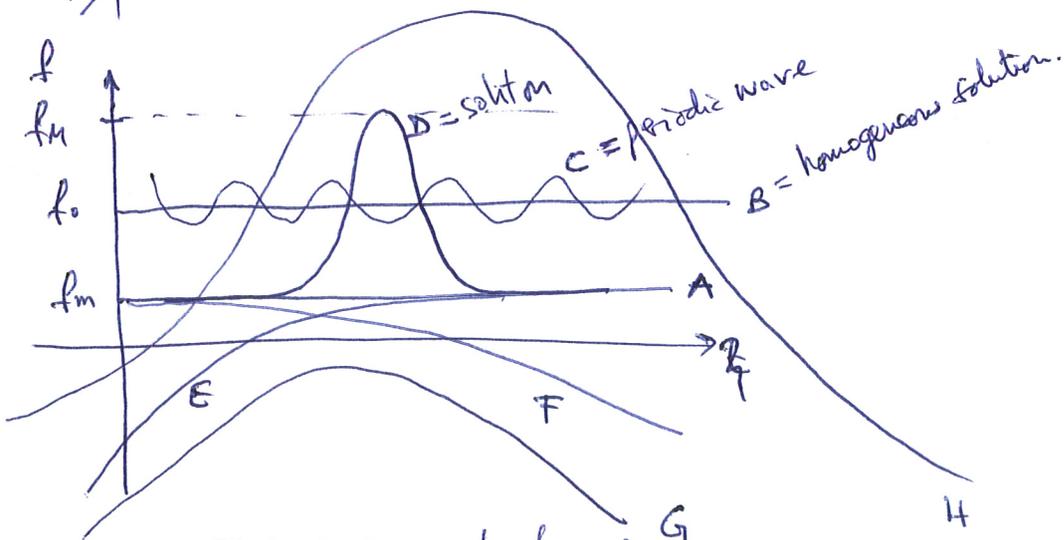


Homoclinic connection
" sech soliton

In general: $v(f) = 2f^3 - cf^2 + Af + B \Rightarrow mf'' = -\frac{\partial v}{\partial f} = -(6f^2 - 2cf + A)$



Use pplane.m to show solutions



E, F, G, H: not physical

Oscillatory solutions of KdV

$$u_t + 6uu_x + u_{xxx} = 0$$

[note sign "+" when compared to Drazin. → This changes signs all over the place]

$$u = f(z) = f(x - ct)$$

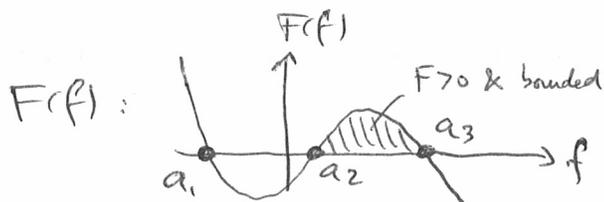
$$\Rightarrow -cf' + 6ff' + f''' = 0$$

$$\int \Rightarrow -cf + 3f^2 + f'' = -A/2 \xrightarrow{\times f'} -cff' + 3f^2f' + f'f'' = -\frac{A}{2}f'$$

$$\int \Rightarrow -\frac{c}{2}(f^2)' + (f^3)' + \frac{1}{2}(f'^2)' = -\frac{A}{2}f' \Rightarrow -\frac{c}{2}f^2 + f^3 + \frac{1}{2}f'^2 = -\frac{A}{2}f - \frac{B}{2}$$

$$\xrightarrow{\times 2} (f')^2 = -2f^3 + cf^2 - Af - B \Rightarrow (f')^2 = f^2(c - 2f) - (Af + B)$$

$$(f')^2 = F(f)$$



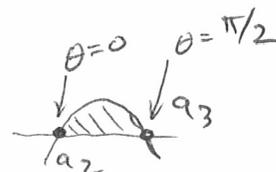
$$\frac{df}{\sqrt{F(f)}} = \pm \int \frac{df}{\sqrt{F(f)}} = \int dz = z - x_0$$

Now, we will apply some change of variables to bring integral to a "familiar" form → elliptic functions.

• let us factorize $F(f) = -2(f - a_1)(f - a_2)(f - a_3)$

note: $c = 2(a_1 + a_2 + a_3)$

• change of variables $f = a_2 + (a_3 - a_2) \sin^2 \theta$



$$df = (a_3 - a_2) 2 \cos \theta \sin \theta d\theta$$

$$\therefore I = \int \frac{df}{\sqrt{F}} = \int \frac{2(a_3 - a_2) \cos \theta \sin \theta d\theta}{\sqrt{-2(f - a_1)(f - a_2)(f - a_3)}} d\theta$$

$$a_2 \frac{c^2}{(1 - s^2)} - a_3 \frac{c^2}{(1 - s^2)}$$

$$(f - a_1)(f - a_2)(f - a_3) = (a_2 + (a_3 - a_2)s^2 - a_1)(a_2 + (a_3 - a_2)s^2 - a_2)(a_2 + (a_3 - a_2)s^2 - a_3)$$

~~$$= (a_3 - a_2) s^2$$~~

$$= (a_2 c^2 + a_3 s^2 - a_1)(a_3 - a_2) s^2 (a_2 - a_3) c^2$$

$$= -(a_2 c^2 + a_3 s^2 - a_1) (a_3 - a_2) s^2 c^2$$

$$= -(a_2 - a_2 s^2 + a_3 s^2 - a_1) (a_3 - a_2)^2 s^2 c^2$$

$$= -(a_2 - a_1 + (a_3 - a_2) s^2) (a_3 - a_2)^2 s^2 c^2$$

$$= -(a_2 - a_1) \left[1 + \frac{a_3 - a_2}{a_2 - a_1} s^2 \right] (a_3 - a_2)^2 s^2 c^2$$

$$b^2 \equiv \frac{a_3 - a_2}{a_2 - a_1}$$

$$\Rightarrow I = \int \frac{2(a_3 - a_2) ds}{\sqrt{2(a_2 - a_1)(1 + b^2 s^2)(a_3 - a_2)^2 s^2 c^2}} d\theta = \sqrt{\frac{2}{a_2 - a_1}} \int \frac{1}{\sqrt{1 + b^2 \sin^2 \theta}} d\theta$$

$$\therefore \pm \sqrt{\frac{z}{a_2 - a_1}} \int_0^\phi \frac{d\theta}{\sqrt{1 - k^2 \sin^2 \theta}} = \zeta - x_0$$

$$\phi : f = a_2 + (a_3 - a_2) \sin^2 \phi$$

$$\Rightarrow \int_0^\phi \frac{d\theta}{\sqrt{\dots}} = \pm \sqrt{\frac{a_2 - a_1}{z}} (\zeta - x_0) =$$

Similar to $\int \frac{1}{\sqrt{1-x^2}} = \sin^{-1} x$

Now: Jacobi elliptic integral: $u = \int_0^\phi \frac{d\theta}{\sqrt{1 - k^2 \sin^2 \theta}} \Leftrightarrow \text{sn}(u, k) = \sin \phi$

$$\Rightarrow \text{sn} \left[\pm \sqrt{\frac{a_2 - a_1}{z}} (\zeta - x_0) \right] = \sin \phi = \sqrt{\frac{f - a_2}{a_3 - a_2}}$$

$$\begin{aligned} \Rightarrow \text{sn}^2 [\dots] &= \frac{f - a_2}{a_3 - a_2} \Rightarrow f = (a_3 - a_2) \text{sn}^2[\dots] + a_2 \\ &= (a_3 - a_2) (1 - \text{cn}^2) + a_2 \\ &= a_3 + (a_3 - a_2) \text{cn}^2 \end{aligned}$$

But: $\text{sn}^2 + \text{cn}^2 = 1$

$$\therefore f(z) = a_3 + (a_3 - a_2) \text{cn}^2 \left[\sqrt{\frac{a_2 - a_1}{z}} (\zeta - x_0) \right]; k$$

$$\Rightarrow u(x, t) = a_3 + (a_3 - a_2) \text{cn}^2 \left[\sqrt{\frac{a_2 - a_1}{z}} (x - ct - x_0) \right]; k$$

where $c = 2(a_1 + a_2 + a_3)$
 $k^2 = (a_3 - a_2) / (a_2 - a_1)$

cn, sn

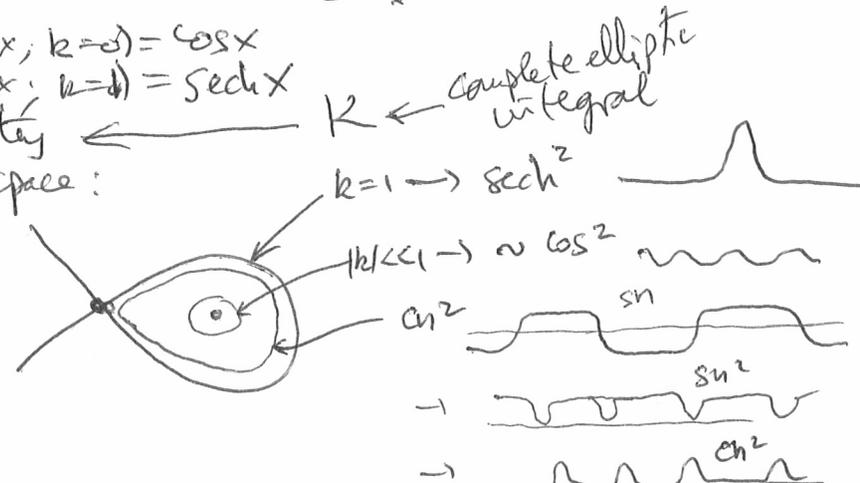
- Jacobi elliptic cosine and sine, pronounced
- cn-wave \rightarrow cnoidal wave

sn: "s"-en
 cn: "c"-en

- properties:
- many many, ... check online
 - cn, sn interpolate between trig. & hyper.
 - k: elliptic modulus [Mittal's use $m = k^2$]

$$\begin{aligned} \text{cn}(x, k=0) &= \cos x \\ \text{sn}(x, k=0) &= \text{sech } x \end{aligned}$$

Remember phase space:



Show cn-m

Kdv-Ellipt. m

Coefficients of cnoidal wave

Since we know that solution to KdV is cnoidal, let us force

$$u(x,t) = A + B \operatorname{cn}^2 [D(x-ct-x_0); k]$$

into KdV and find restrictions/relations between params (A, B, D, c, k)

• We'll need diff. flas. for cn & sn: $\begin{aligned} \operatorname{sn}' &= \operatorname{cn} \operatorname{dn} \\ \operatorname{cn}' &= -\operatorname{sn} \operatorname{dn} \\ \operatorname{dn}' &= -k^2 \operatorname{cn} \operatorname{sn} \end{aligned}$

$$\therefore \text{KdV: } u_t + 6u u_x + u u_{xx} = 0$$

$$\Rightarrow -c D B^2 (-\operatorname{sn}) \operatorname{dn} / \operatorname{cn} + 6(A + B \operatorname{cn}^2)(B D (-\operatorname{sn}) \operatorname{dn} \operatorname{cn}) + [B D (-\operatorname{sn}) \operatorname{dn} (\operatorname{cn})]_{xx} = 0$$

... gets messy.

let us use Maple

KdV-
elliptic-
sol.mws

$$\dots \Rightarrow \begin{cases} B = 2D^2 k^2 \\ c = 6A + 6B = 4D^2(1+k^2) \end{cases}$$

Ex: $A=0$ & $k=1$ should recover sech soln:

$$\begin{aligned} \begin{cases} B = 2D^2 \\ c = 6B = 8D^2 \end{cases} &\Rightarrow c = 12D^2 - 8D^2 = 4D^2 \\ &\Rightarrow D = \frac{\sqrt{c}}{2} \Rightarrow B = 2D^2 = \frac{c}{2} \end{aligned}$$

$\therefore A=0, B=c/2, D=\sqrt{c}/2, k=0$ gives:

$$\boxed{u(x,t) = \frac{c}{2} \operatorname{sech}^2 \left[\frac{\sqrt{c}}{2} (x-ct-x_0) \right]} \quad \checkmark$$

Other solutions (not traveling with const. shape)

- (A) ~~Similarity solutions~~
- (B) ~~Rational solutions~~

(A) Similarity solutions: suppose $u(x,t) = t^m f(\eta)$ and $\eta = x t^n$
 \rightarrow x space and time linked

$u = -(3t)^{-2/3} f(\eta)$, $\eta = x \cdot (3t)^{-1/3}$ [see Brazi sec 2.6]
 $\Rightarrow f''' + (4-\eta)f'' - 2f' = 0 \leftarrow \text{LINEAR!}$
 [c=0 \Rightarrow stationary]

(B) Rational solutions: assume $u(x,t) = f(x)$
 $\Rightarrow 6\eta f' + f''' = 0$

$\int d\eta \Rightarrow 3(f')^2 + f'' = A$

Suppose u & derivative $\xrightarrow{x \rightarrow \pm\infty} 0 \Rightarrow A = 0 \Rightarrow 3f'^2 + f'' = 0$ (1)

$x f' \Rightarrow 3 f' f'^2 + f' f'' = 0$

$\int dx \Rightarrow (f')^3 + \frac{1}{2}(f'')^2 = 0$

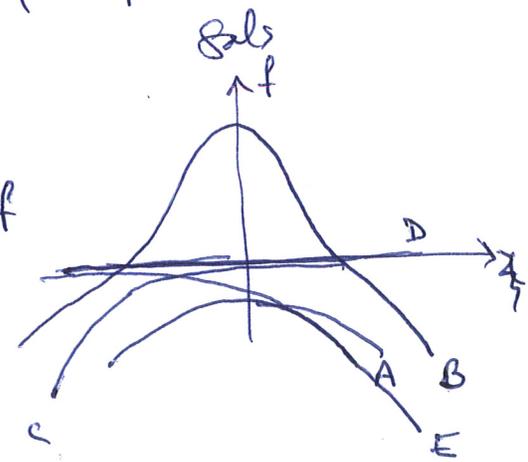
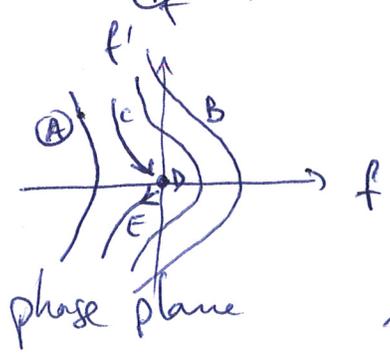
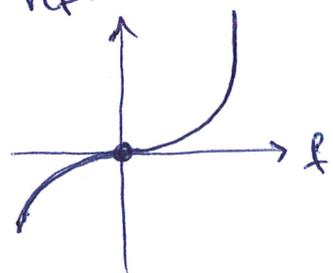
$\Rightarrow f' = \pm \sqrt{-2f^3}$

$\Rightarrow \int \frac{df}{(-f)^{3/2}} = \pm \sqrt{2} \int dx \Rightarrow \frac{(-f)^{-1/2}}{-1/2} = \pm \sqrt{2}(x - x_0)$

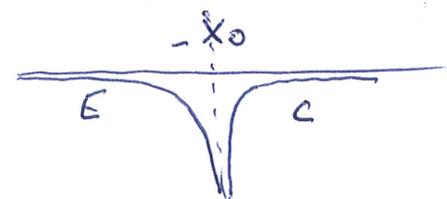
$\Rightarrow -f = \frac{2}{(x-x_0)^2} \Rightarrow \boxed{f = \frac{-2}{(x-x_0)^2}}$ (2)

Where does this sol. come from?

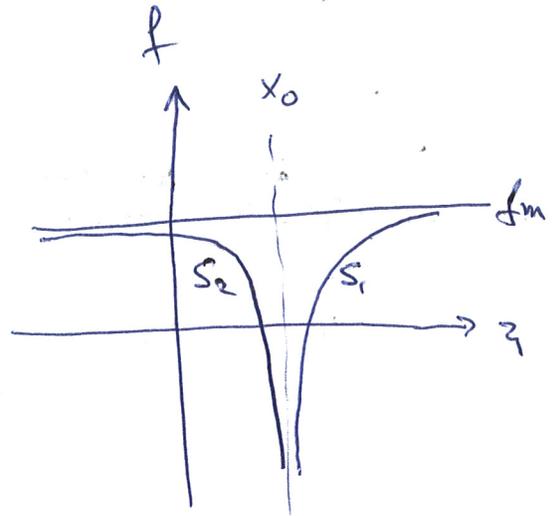
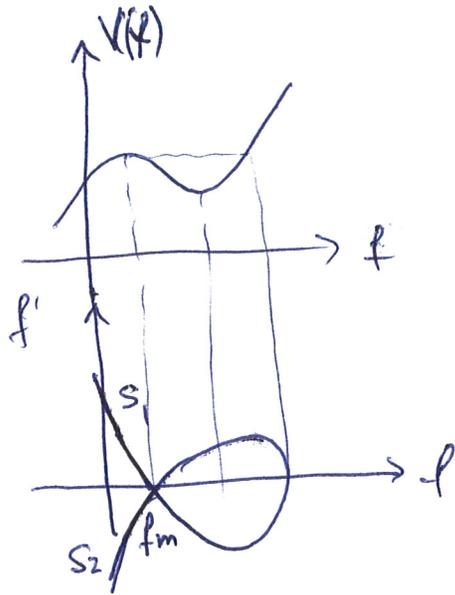
(1) $\Rightarrow f'' = -3f^2 = -\frac{dV}{df}$, $V(f) = f^3$



Solution (2) is nothing that the concatenation of E and C



there is a whole family of rational solutions when A & B const. are left:



Q: how can we obtain these anal. solutions using the $c=0$ solution $f(z) = \frac{-z}{(z-x_0)^2}$?

A: by Galilean invariance! \rightarrow change background and velocity.

- these rational solutions are not physical due to their singularity.
- there are other rational solutions not included in above description.

Breathing Solutions (Carbion)

Take the modified KdV eq = mKdV with $N=2$ ($n=1 \Rightarrow$ standard KdV)

$$u_t + (Nn)(\phi(z)) u^N u_x + U_{xxx} = 0$$

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 $u_t + u^2 u_x + U_{xxx} = 0$
 \rightarrow rescale $u = u\sqrt{2}$
 \Rightarrow

$$u_t + 6u^2 u_x + U_{xxx} = 0 \quad \text{transform using } u = U_x \text{ \& } \phi = \tan(\frac{1}{2}U)$$

$$\Rightarrow (1 + \phi^2)(\phi_t + \phi_{xxx}) + 6\phi_x(\phi_x^2 - \phi_{xx}) = 0$$

$$\Rightarrow \text{same trick as in HW7}$$

The following is a solution:

$$u(x,t) = -2 \frac{\partial}{\partial x} \left[\tan^{-1} \left(\frac{l}{k} \underbrace{\sin(kx + mt + a)}_{\text{traveling sin}} \underbrace{\text{sech}(l(x + nt + b))}_{\text{traveling sech}} \right) \right]$$

$k, l, a, b \in \mathbb{R}$
with $m = k(k^2 - 3l^2)$ &
 $n = l(3k^2 - l^2)$

show
mKdV-breather-m

two soliton EXACT solution types of sds

→ more on how to obtain these later.

KdV: $u_t + 6uu_x + u_{xxx} = 0$

⇒ 2 soliton solution

$$u(x,t) = 2 \frac{k_1^2 E_1 + k_2^2 E_2 + 2(k_2 - k_1)^2 E_1 E_2 + A_2 (k_1^2 E_1 + k_2^2 E_2) E_1 E_2}{(1 + E_1 + E_2 + A_2 E_1 E_2)^2}$$

where $E_i = e^{(k_i x - k_i^3 t + \alpha_i)}$, $k_i, \alpha_i \in \mathbb{R}$; $A_2 = \frac{(k_2 - k_1)^2}{(k_1 + k_2)^2}$

In particular: $u(x,t) = -12 \frac{3 + 4 \cosh(2x - 8t) + \cosh(4x - 64t)}{[3 \cosh(x - 28t) + \cosh(3x - 36t)]^2}$

$t \rightarrow -\infty$: $u(x,t) \rightarrow +8 \operatorname{sech}^2 \left[\frac{1}{2}(x - 16t) + \frac{1}{4} \log 3 \right] + 2 \operatorname{sech}^2 \left[x - 4t + \frac{1}{2} \log 3 \right]$

$t \rightarrow +\infty$: $u(x,t) \rightarrow$ - +

∴ the fast soliton ($v=16$) is advanced by ~~2x~~ $2 \times \frac{1}{4} \log 3 = \frac{1}{2} \ln 3$
and the slow soliton ($v=4$) is retarded by $2 \times \frac{1}{2} \log 3 = \ln 3$

Shree Mattar
KdV - 2 sol.m

→ show flux for more solitons from [Shen]
Conservation laws for KdV

Ex: if $\rho(x,t)$ = density of gas
 $u(x,t)$ = velocity of gas (in x-direction)

continuity eq. ⇒ $\frac{\partial \rho}{\partial t} + \frac{\partial}{\partial x} (\rho u) = 0$ (1)

If gas is confined ⇒ $\rho(x=\pm\infty) = 0$

⇒ $\int_{-\infty}^{\infty} \rho dx \Rightarrow \frac{d}{dt} \int_{-\infty}^{\infty} \rho dx + [\rho u]_{-\infty}^{+\infty} = 0$

∫_{-∞}[∞] ρ dx = const, $\rho_t + (\rho u)_x = 0$
conservation of mass!

continuity eq:

$\rho(x,t)$
dx

change $dm_1 = \rho_t dt dx$

u
x x+dx

change $dm_2 = \rho(x+dx)u(x,t) - \rho(x)u(x,t) dt$
→ $u(x) \rho_t dx$

$dm_1 + dm_2 = 0$

⇒ $\rho_t dx + [\rho(x+dx)u(x,t) - \rho(x)u(x,t)] dt = 0$

⇒ $\rho_t + \frac{d(\rho u)}{dx} = 0$

KdV: let us construct conservation laws for KdV, ie. $\frac{d}{dt} F(u, u_x, u_t, u_{xx}, \dots) = 0$

$$u_t + 6uu_x + u_{xxx} = 0$$

$$\Rightarrow \frac{\partial u}{\partial t} + \frac{\partial}{\partial x} [3u^2 + u_{xx}] = 0$$

\therefore use same trick as for the gas ^{flux}

$$\int dx \Rightarrow \frac{d}{dt} \int_{-\infty}^{\infty} u(x,t) dx + [3u^2 + u_{xx}]_{-\infty}^{\infty} = 0$$

\therefore if solution decays @ $\pm\infty$, $\lim_{x \rightarrow \pm\infty} u(x,t) = 0$

$$\Rightarrow \boxed{\int_{-\infty}^{\infty} u(x,t) dx = \text{const.}} \quad \text{conservation of mass (water)}$$

Δ if $u(x,t)$ is ~~not~~ periodic on x then the integral for mass can be done on a single period (otherwise $\int_{-\infty}^{\infty} = \infty$)

More const. of motion: massage KdV and integrate.

$$\text{KdV: } u_t + 6uu_x + u_{xxx} = 0$$

$$\times u \quad u u_t + 6u^2 u_x + u u_{xxx} = 0$$

parts: $\int f g' = \int [(f g)' - f' g] = \int [(u u_{xx})_x - u_x u_{xx}] = u u_{xx} - \frac{u_x^2}{2}$

$$\int \Rightarrow \int \left[\frac{(u^2)_t}{2} + \left[2(u^3)_x + u u_{xx} - \frac{u_x^2}{2} \right] \right] dx = 0$$

$$\Rightarrow \frac{1}{2} \frac{d}{dt} \int_{-\infty}^{\infty} u^2(x,t) dx + [2u^3 + u u_{xx} - \frac{u_x^2}{2}]_{-\infty}^{\infty} = 0$$

if u decays to zero if not, just subtract the constant background.
 \leftrightarrow regularizing (use δ 's)

$$\therefore \boxed{\int_{-\infty}^{\infty} u^2(x,t) dx = \text{const}} \quad \text{conservation of momentum.}$$

physics: $\underline{p} = m v$, here $v \propto u$ and $u \propto$ mass

momentum $\therefore p \propto u^2$ because $u_t + 6uu_x + u_{xxx} = 0$

Another: Energy: $3u^2(\text{KdV}) + u_x \frac{\partial}{\partial x}(\text{KdV}) = 0 \Rightarrow$ (HW)

$$\boxed{\int_{-\infty}^{\infty} \left(u^3 + \frac{1}{2} u_x^2 \right) dx = \text{const}} \quad \text{conservation of Energy}$$

\uparrow potential \uparrow kinetic

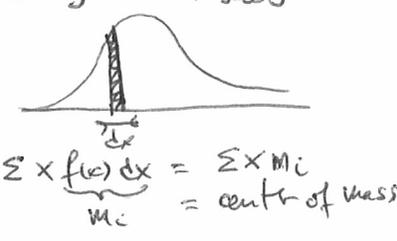
It is possible to prove that KdV has an ∞ # of independent constants of motion / conservation laws

- Conservation laws very useful:
- * project solutions onto trial functions
 - * check that numerics are doing ok.

1/2/08

Center of Mass (not a const. of motion)

- Center of mass of distribution $f(x)$: $x_c = \int_{-\infty}^{\infty} x f(x) dx / \int_{-\infty}^{\infty} f(x) dx$
- for us: $x_c(t) = \int_{-\infty}^{\infty} x u(x,t) dx / \int_{-\infty}^{\infty} u dx$
- let us follow x_c for KdV:



$$u_t + 6u u_x + u_{xxx} = 0$$

$$\Rightarrow u_t + 3(u^2)_x + u_{xxx} = 0$$

$$\Rightarrow u_t + \partial_x [3u^2 + u_{xx}] = 0$$

$$\Rightarrow \partial_t \int_{-\infty}^{\infty} x u dx + \partial_x \int_{-\infty}^{\infty} x [3u^2 + u_{xx}] dx - \int_{-\infty}^{\infty} [3u^2 + u_{xx}] dx = 0$$

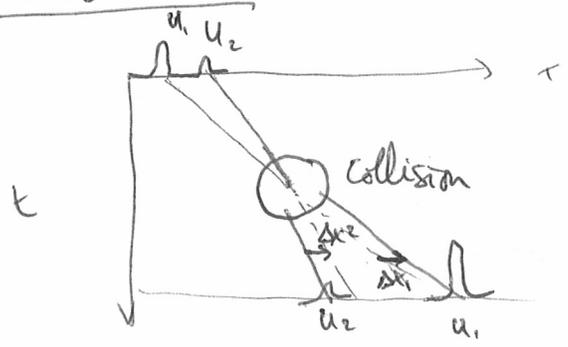
$(uv)' = u'v + uv'$
 $uv' = (uv)' - u'v$

If u decays fast enough $xu^2, xu_{xxx} \xrightarrow{x \rightarrow \pm \infty} 0 \Rightarrow \frac{d}{dt} \int x u dx = K_1 + 3 \int u^2$
 $\Rightarrow \frac{d}{dt} \left[\frac{\int x u}{\int u} \right] = K_2$
 $\rightarrow \frac{d}{dt} [x_c(t)] = K_2 \therefore x_c(t) = K_2 t + K_3$
 $\therefore \boxed{x_c(t) = v_c t + x_0} \leftarrow \text{COM moves @ const. speed!}$

2/27/00

Application to soliton collisions

2-soliton solution:



Before collision:

$$u_1 \approx \frac{v_1}{2} \operatorname{sech}^2 \left[\frac{\sqrt{v_1}}{2} (x - v_1 t - x_1) \right]$$

$$u_2 \approx \frac{v_2}{2} \operatorname{sech}^2 \left[\frac{\sqrt{v_2}}{2} (x - v_2 t - x_2) \right]$$

After collision:

suppose u_1 is displaced by Δx_1 :

$$u_1 \approx \frac{v_1}{2} \operatorname{sech}^2 \left[\frac{\sqrt{v_1}}{2} (x - v_1 t - x_1 - \Delta x_1) \right]$$

\Rightarrow how much is u_2 displaced? $\propto \Delta x_2 = \Delta x_2(\Delta x_1)$

$$u_2 \approx \frac{v_2}{2} \operatorname{sech}^2 \left[\frac{\sqrt{v_2}}{2} (x - v_2 t - x_2 - \Delta x_2) \right]$$

center of mass:

before: $\int x u dx = At + B \Rightarrow \int x u_1 + \int x u_2 = At + B \quad \text{① } (u \approx u_1 + u_2)$

but: $\int_{-\infty}^{\infty} x \operatorname{sech}^2 a(x-b) dx \stackrel{\text{Maple}}{\approx} \frac{M_{\text{pole}}}{a^2} \frac{2b}{a}$ show assume trick in Maple

$$\therefore \int x u_1 = \frac{v_1}{2} \int 2 (v_1 t + x_1) \cdot \frac{2}{\sqrt{v_1}} = 2\sqrt{v_1} (v_1 t + x_1)$$

① $\Rightarrow 2\sqrt{v_1} (v_1 t + x_1) + 2\sqrt{v_2} (v_2 t + x_2) = At + B$

After: $2\sqrt{v_1} (v_1 t + x_1 + \Delta x_1) + 2\sqrt{v_2} (v_2 t + x_2 + \Delta x_2) = At + B$

$$\ominus -\cancel{2\sqrt{v_1}} \Delta x_1 - \cancel{2\sqrt{v_2}} \Delta x_2 = 0 \Rightarrow \boxed{\frac{\Delta x_2}{\Delta x_1} = -\sqrt{\frac{v_1}{v_2}}}$$

For the 2-sol. solution that we saw:

$$v_1 = 16, \Delta x_1 = \frac{1}{2} \ln 3$$

$$v_2 = 4, \Delta x_2 = -\sqrt{\frac{v_1}{v_2}} \Delta x_1 = -\sqrt{4} \frac{1}{2} \ln 3 = -\ln 3 \quad \checkmark$$

Δ center of mass need to move at const. speed even DURING collision.

5.1.2 More on Conservation Laws: $\infty \neq !$

KdV: $u_t - 6uu_x + u_{xxx} = 0$ [note I am using "-" as in Drazni's book]

+ Miura transf: $u = v^2 + v_x$
 + another transf: $v = \frac{1}{2} \epsilon^{-1} + \epsilon w$ $\epsilon \in \mathbb{R}$

$\Rightarrow u = \frac{1}{4} \epsilon^{-2} + w + \epsilon^2 w^2 + \epsilon w_x$

eliminate since we can always add through Galilean invariance

$\therefore \boxed{u = w + \epsilon^2 w^2 + \epsilon w_x}$ (Gardner transf.)

KdV $\xrightarrow[\dots]{\text{Gardner}}$ $(1 + \epsilon \frac{\partial}{\partial x} + 2\epsilon^2 w) [w_t - 6(w + \epsilon^2 w^2)w_x + w_{xxx}] = 0$

$\therefore u$ KdV sol $\Rightarrow w$ sol of $v = 0$

$\epsilon = 0 \Rightarrow w = u$ and recover KdV

+ Now search for consv. quant: $\epsilon \neq 0$

$\frac{\partial}{\partial t}(w) + \frac{\partial}{\partial x}(-3w^2 + 2\epsilon^2 w^3 + w_{xx}) = 0$

$\int_{-\infty}^{\infty} dx \Rightarrow \frac{d}{dt} \int_{-\infty}^{\infty} w dx + \left[\dots \right]_{-\infty}^{\infty} = 0$
 decay for w .

$\therefore \int_{-\infty}^{\infty} w dx = \text{const of motion.}$

+ since $w \xrightarrow{\epsilon \rightarrow 0} u$ we can expand w by

$w = \sum_{n=0}^{\infty} \epsilon^n w_n$ when $|\epsilon| \ll 1$

and we need $\int w dx = \text{const} \Rightarrow \int w_n dx = \text{const}$ since ϵ is arbitrary.

use Gardner transf: $u = w + \epsilon^2 w^2 + \epsilon w_x$

+ expansion $u = \sum \epsilon^n w_n + \epsilon^2 \left(\sum \epsilon^m w_m \right)^2 + \epsilon \sum \epsilon^n w_{nx}$

$\Rightarrow \sum \epsilon^n w_n = u - \sum \epsilon^{n+1} w_{nx} - \epsilon^2 (\epsilon^0 w_0 + \epsilon^1 w_1 + \epsilon^2 w_2 + \dots)^2$

$\Rightarrow \sum \epsilon^n w_n = u - \sum \epsilon^{n+1} w_{nx} - \epsilon^2 (w_0^2 + \epsilon^2 w_1^2 + \dots + 2\epsilon w_0 w_1 + 2\epsilon^3 w_1 w_2 + \dots + 2\epsilon^2 w_0 w_2 + \dots)$

$\epsilon^0: w_0 = u$

$\epsilon^1: w_1 = -w_{0x}$

$\epsilon^2: w_2 = -w_{1x} - w_0^2$

$\epsilon^3: w_3 = -w_{2x} - 2w_0 w_1$

$\epsilon^4: w_4 = -w_{3x} - w_1^2 - 2w_0 w_2$

;

& since each $\int u_n = \text{const}$ we can solve progressively 53

$$\epsilon^0: \int u = ct \quad (\text{Mass})$$

$$\epsilon^1: \int u_x = ct$$

$$\epsilon^2: \int + u_{xx} - u^2 = ct \quad (\text{Energy})$$

$$\epsilon^3: \int - (u_{xx} - u^2)_x - 2u(-u_x) = ct$$

$$\epsilon^4: \int [-(u_{xx} - u^2)_x + 2uu_x]_x - (u_x)^2 - 2u(u_{xx} - u^2) = ct$$

Note that all ϵ^{odd} are made up of exact differentials.

$$\epsilon^1: \int_{-\infty}^{\infty} u_x dx = [u]_{-\infty}^{\infty} = 0$$

$$\epsilon^{\text{odd}}: \int_{-\infty}^{\infty} \frac{d}{dx} [f(u)] dx = [f(u)]_{-\infty}^{\infty} = 0$$

$\therefore \epsilon^{\text{odd}}$ do NOT give any conservation laws

u decays fast enough.

the others:

$$\epsilon^2: \int \frac{d}{dx} [u_x] - u^2 = ct \Rightarrow [u_x]_{-\infty}^{\infty} - \int u^2 = ct \quad (\text{Momentum})$$

$$\epsilon^4: \int [-(u_{xx})_x + 2uu_x]_x - (u_x)^2 - 2u(u_{xx} - u^2) = ct$$

$$\Rightarrow \int u_x^2 - 2 \frac{d}{dx} (uu_x) + 2u_x^2 + 2u^3 = ct$$

$$\Rightarrow [uu_x]_{-\infty}^{\infty} + \int u_x^2 + 2u^3 = ct$$

$$\Rightarrow \int \frac{u_x^2}{2} + u^3 = ct \quad (\text{Energy})$$

It can be shown that ϵ^{2k} has terms up to u^{k+1} and thus indep. from lower order $[\epsilon^{2(k-1)}]$ terms $\Rightarrow \infty \#$ of conservation laws !!!

Recurrence for conservation laws

[D: p94-95]

$$P^0 = u$$

$$P^n = u + \epsilon^2 [P^{n-1}]^2$$

$$\int_{-\infty}^{\infty} P^n dx = ct$$

$$n=0: \int u = ct$$

$$n=1: P^1 = u + \epsilon^2 [P^0]^2 = u + \epsilon^2 [u]^2$$

$$\therefore \int u + \epsilon^2 \int u^2 = ct \Rightarrow \int u^2 = ct$$

$$n=2: P^2 = u + \epsilon^2 [P^1]^2 = u + \epsilon^2 [u + \epsilon^2 u^2]^2$$

$$= u + \epsilon^2 u^2 + 2\epsilon^4 u^3 + \epsilon^6 u^4$$

$$\therefore \int u + \epsilon^2 \int u^2 + \epsilon^4 \int 2u^3 + \epsilon^6 \int \dots$$

5.4 Bäcklund Transformation (BT)

consider 2 uncoupled, nonlinear, PDEs:

$$\begin{cases} P(u) = 0 \\ Q(v) = 0 \end{cases}$$

i.e. $\forall u, v$ s.t. $P(u) = 0$ & $Q(v) = 0$
 If u, v such that $R_i = 0 \iff \begin{cases} P(u) = 0 \\ Q(v) = 0 \end{cases}$

let R_1 & R_2 : $R_i(u, v, u_x, v_x, u_y, v_y, \dots; x, t) = 0$
 be 2 relations such that if R_i is integrable for v when $P(u) = 0$
 and v is a sol. to $Q(v) = 0$ and vice-versa $\implies R_i = 0$

is the Bäcklund transf. of Eqs $P(u)$ & $Q(v)$.
 E.g. If u sol. to $P(u) = 0 \implies v$ is a sol. to $Q(v) = 0$

If the 2 eqs are same: $P = Q$ then $R_i = 0$ are called auto-Bäcklund transf.

Use: find one solution auto find another!

Ex: consider $R_1: u_x = v_y$
 $R_2: u_y = -v_x$

Bäcklund?: u sol: $u_{xx} + u_{yy} = 0$
 $R_1, R_2 \implies u_x + v_y = 0$
 then $v_{xx} + v_{yy} = (-u_x)_x + (u_x)_y$
 $= 0 \checkmark$
 \therefore yes it is a BT-
auto

and Laplace eq: $u_{xx} + u_{yy} = 0$
 $v_{xx} + v_{yy} = 0$

one solution $v(x, y) = xy$
 $\implies \begin{cases} u_x = v_y = x \implies u = \frac{x^2}{2} + f(y) \\ u_y = -v_x = -y \implies u = -\frac{y^2}{2} + g(x) \end{cases} \implies u = \frac{1}{2}(x^2 - y^2)$
 $u_{xx} + u_{yy} = 0 \checkmark$
 is another solution to Laplace Eq.

5.4.2 KdV: use KdV and Miura transf:

KdV: $u_t - 6uu_x + u_{xxx} = 0$ (1)

Miura transf: $u = v^2 + v_x$ (Miura+)
 gal - u: $u = \lambda + v^2 + v_x$ (2) (Galilean invariance)

\implies Modified KdV: $[u_t - 6(v^2 + \lambda)v_x + v_{xxx}] = 0$ (3)

(3) \iff if v sol to Miura $\iff u$ sol to KdV

Note that if v sol to (3) then $-v$ is also sol

\therefore take (2): $\begin{cases} u_+ = \lambda + v^2 + v_x \\ u_- = \lambda + v^2 - v_x \end{cases}$

$\implies u_+ + u_- = 2(\lambda + v^2); u_+ - u_- = 2v_x$

transf: $u_{\pm} = (w_{\pm})_x \implies \begin{cases} (w_+ + w_-)_x = 2\lambda \\ w_+ - w_- = 2v + c_2 \end{cases}$

o transf: $U_{\pm} = (W_{\pm})_x$ ^(3.5) $\Rightarrow \begin{cases} W_+ - W_- = 2U + C_2 \\ \text{and } (W_+ + W_-)_x = 2\lambda + 2U^2 \end{cases}$

$\Rightarrow (W_+ + W_-)_x = 2\lambda + \frac{1}{2}(W_+ - W_- - C_2)^2$

55
32 4/5

choose $C_2 = 0$ $\begin{cases} W_+ - W_- = 2U \\ (W_+ + W_-)_x = 2\lambda + \frac{1}{2}(W_+ - W_-)^2 \end{cases}$ (4)

o transf Modified KdV using $(U_{\pm}) = (W_{\pm})_x$ of (4)

(3) with (3)&(4) $\Rightarrow (W_+ - W_-)_t - 3(W_+^2 - W_-^2)_x + (W_+ - W_-)_{xxx} = 0$ (5)

(4b) and (5) are the auto-Bäcklund transf. of the mKdV (5) [(4b) is the "x" transf. & (5) is the "t" transf.]

o Fixed sols of KdV

- give a 1st sol: trivial sol: $W = 0$
- use $W_+ = 0$ ^(4b) $\Rightarrow (W_+ + 0)_x = 2\lambda + \frac{1}{2}(W_+ - 0)^2$

$u=0 \Rightarrow W_x=0 \Rightarrow W = \text{const}$
10/21/08

$\Rightarrow W_x = 2 \times \frac{1}{2} (W_+)^2$
 ~~$W_x = 2 \times \frac{1}{2} (W_+)^2$~~

$\int dx$

$W_{+x} = 2\lambda + \frac{1}{2}W_+^2$ (6)

$W_+(x,t) = -2k \tanh[kx + f(t)]$

$k = \sqrt{-\lambda}$

$\left[\begin{array}{l} |W_+| < 2k \\ \text{if } |W_+| > 2k \\ \Rightarrow \tanh \rightarrow \coth \end{array} \right]$

- for $f(t)$ we need (5) to be satisfied:

plug $W_+ = \tanh$ into (5) and solve for $f(t)$

$W_{+t} - 3W_{+x}^2 + W_{+xxx} = 0$

but $W_{+xxx} = \frac{\partial}{\partial x} (W_{+xx}) = \frac{\partial}{\partial x} (W_+ W_{+x}) = W_{+x} W_{+xx} + W_+ W_{+xxx}$

$\Rightarrow W_{+t} - 3W_{+x}^2 + W_{+x}^2 + W_+ W_{+xxx} = 0$

$\Rightarrow W_{+t} + 2W_{+x} (W_{+x} - \frac{1}{2}W_{+xxx}) = 0$

$\Rightarrow f(t) = -4k^3 t - kx_0$ $k = \sqrt{-\lambda}$

$\therefore W_+(x,t) = -2k \tanh[k(x-x_0 - 4k^2 t)]$

is another solution to mKdV.

- remember $U_+ = (W_+)_x$ ~~$U_+ = (W_+)_x$~~ $(\tanh' = \text{sech}^2)$

$\Rightarrow \boxed{U_+ = -2k^2 \text{sech}^2[k(x-x_0 - 4k^2 t)]}$ soliton sol. of KdV

\therefore vacuum $\xrightarrow{\text{Bäcklund BT}}$ 1 soliton solution \uparrow vel = $4k^2$

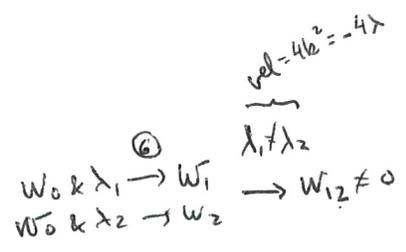
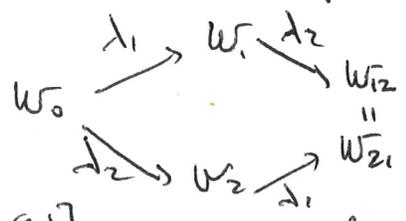
This process could be applied as many times as we want to produce more solutions. Everytime we do this we need one x-integration and one t-integration → there is a more elegant way:

5.4.3 the RLV Bäcklund transf: an algebraic relation

- Suppose that one starts with a solution w_0 and one builds two solutions using the BT for $2 \neq \lambda$'s: $\lambda_1 \rightarrow w_1, \lambda_2 \rightarrow w_2$
- one can write:

$$\begin{cases} (w_1 + w_0)_x = 2\lambda_1 + \frac{1}{2}(w_1 - w_0)^2 & (6) \\ (w_2 + w_0)_x = 2\lambda_2 + \frac{1}{2}(w_2 - w_0)^2 & (6) \end{cases}$$
- now construct another 2 solutions: w_{12} from w_1 and λ_2, w_{21} from w_2 and λ_1 ,

$$\begin{cases} (w_{12} + w_1)_x = 2\lambda_2 + \frac{1}{2}(w_{12} - w_1)^2 & (7) \\ (w_{21} + w_2)_x = 2\lambda_1 + \frac{1}{2}(w_{21} - w_2)^2 & (7) \end{cases}$$
- It turns out, Branchi's theo of permutability, that $w_{12} = w_{21}$ this is because



• Do $[(7)-(7b)] - [(6a)-(6b)]$ and solve for w_{12} :

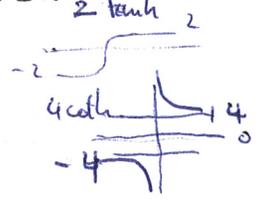
$$w_{12} = w_0 - \frac{4(\lambda_1 - \lambda_2)}{w_1 - w_2}$$

this gives a sort of nonlinear superposition principle if one takes w_0 and λ_1 and λ_2 to be the desired velocities $w_0 = 0$ & $vel_1 = 4b_1^2, vel_2 = 4b_2^2$

- Ex: $vel_1 = 4 \Rightarrow k_1 = 1 \Rightarrow \lambda_1 = -1$
 $vel_2 = -4 \Rightarrow k_2 = 2 \Rightarrow \lambda_2 = -4$

Remember (6) $\Rightarrow w_{\frac{1}{2}x} = 2\lambda + \frac{1}{2}w_{\frac{1}{2}}^2$
 $\Rightarrow \begin{cases} w_1 = -2k \tanh k(x - x_0 - 4k^2t) & \text{if } |w_1| < 2k \\ w_2 = -2k \coth k(x - x_0 - 4k^2t) & \text{if } |w_2| > 2k \end{cases}$

\therefore take $\begin{cases} w_1 = -2 \tanh(x - 4t) & k_1 = 1 \leftarrow \text{non-singular} \\ w_2 = -4 \coth(2x - 32t) & k_2 = 2 \leftarrow \text{singular} \end{cases}$



$$\Rightarrow w_{12} = 0 - \frac{4(1+4)}{-2+4} = \frac{6}{2 \coth - \tanh}$$

and $u_{12} = (w_{12})_x =$ the 2-sol solution that we have \leftarrow non-singular. written a few times. (p25 in my notes)